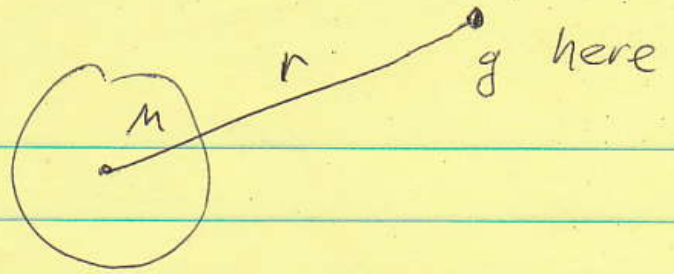


$$g = \frac{GM}{r^2}$$

↑  
gravitational field strength



Near earth's surface:



$$g = 9.8 \text{ m/s}^2$$

$$6.378 \times 10^6 \text{ m}$$

$$R_e = \cancel{1.49 \times 10^8 \text{ km}}$$

$$g = \frac{GM_e}{R_e^2}$$

↑  
Newton

~~$g = \frac{GM_e}{R_e^2}$~~

$G = ? \quad M_e = ?$

Cavendish found G.

$$G = 6.67 \times 10^{-11} \frac{\text{m}^3}{\text{kg} \cdot \text{s}^2}$$

$$M_e = \frac{g R_e^2}{G}$$

$$g = \frac{G M_e}{R_e^2} \rightarrow \frac{\text{m}}{\text{s}^2} = \frac{(\text{?}) \text{ kg}}{\text{m}^2}$$

$$\frac{\text{J} \cdot \text{m}}{\text{kg}^2} = \frac{\text{kg} \cdot \frac{\text{m}^3}{\text{kg} \cdot \text{s}^2}}{\text{kg}^2} = \frac{\text{m}^3}{\text{kg} \cdot \text{s}^2} = (\text{?})$$

$$\hookrightarrow = \frac{\text{J} \cdot \text{m}}{\text{kg}^2}$$

$$G = 6.67 \times 10^{-11} \frac{\text{J} \cdot \text{m}}{\text{kg}^2}$$

Energy is in J.  $G$  is in  $\frac{\text{J} \cdot \text{m}}{\text{kg}^2}$

$$G \propto \frac{U \cdot r}{m_1 m_2}$$

← energy from gravity  
← distance  
← masses

$$U \propto \frac{G m_1 m_2}{r}$$

$\propto$  means  
"is proportional to"

$$U = - \frac{G m_1 m_2}{r}$$

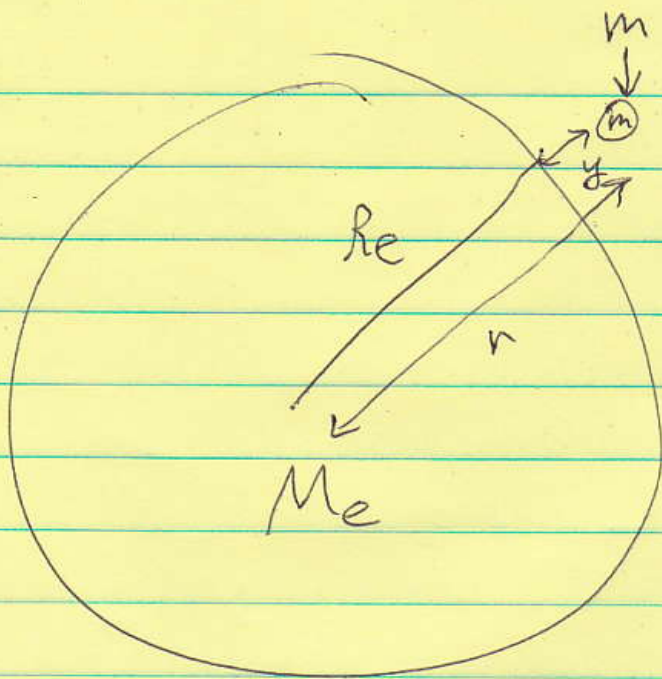
↑  
gravitational potential energy  
caused by masses  $m_1$  &  $m_2$   
interacting gravitationally.

Near earth surface:

$$U = mgy \quad (\text{looks very different})$$

$$g = \frac{G M_e}{R_e^2}$$

$$U = \frac{G M_e m y}{R_e^2}$$



~~$$U = -\frac{GM_em}{r}$$~~

$$r = R_e + y \approx R_e$$

Compare to:

$$U = mgy = \frac{GM_emy}{R_e^2}$$

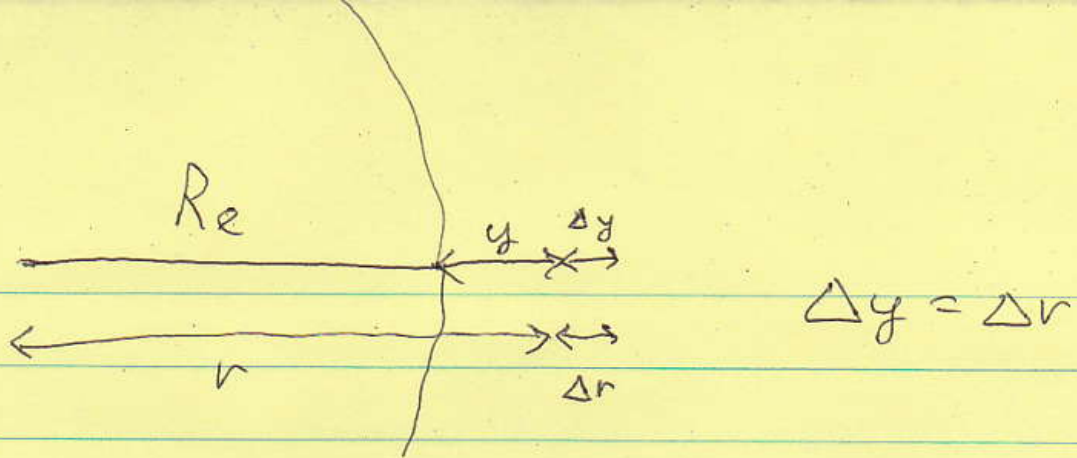
$U$  depends on a coordinate system  
Only changes in  $U$  are real.

We only need formulas for  $\Delta U$   
that agree.

$$\Delta U = \left[ -GM_em \Delta \left( \frac{1}{r} \right) \right] \quad r = R_e + y$$

compare to:

$$\Delta U = mgy = \left[ \frac{GM_em}{R_e^2} \Delta y \right]$$



$$\Delta U = -G M_e m \Delta \left( \frac{1}{r} \right) \approx G M_e m \left( \frac{\Delta r}{r^2} \right)$$

Compare to:

$$\Delta U = m g \Delta y = \frac{G M_e m}{R_e^2} \Delta r$$

$$g = \frac{G M_e}{R_e^2}$$

$$\Delta \left( \frac{1}{r} \right) \approx \frac{-\Delta r}{r^2}$$

For  $\Delta r$  small

Near earth's surface:  $r \approx R_e$

$$G M_e m \left( \frac{\Delta r}{R_e^2} \right) \approx$$

Newton was right;  $\Delta U = m g \Delta y$   
is just an approximation

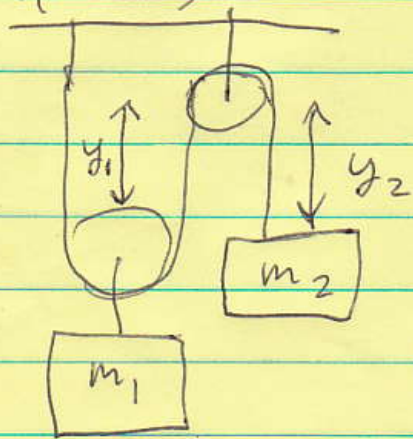
#6 ~~The~~ The ball has maximum kinetic energy at the bottom; at the top, it has converted all its kinetic energy to ~~reach its~~ potential energy, which is highest at the top.

(Anya's to Ivan's) penny)

#12 ~~At the top~~ The ratio of initial energies is 2 because the initial energy  $E$  of each is  $mgy$ , and the heights' ratio is 2. By conservation of energy, the ratio of final energies is also 2. The final (all kinetic) energy  $E$  of each is  $\frac{1}{2}mv^2$ , so  $v = \sqrt{2E/m}$ , so  $v \propto \sqrt{E}$ , so the ratio of final speeds is  $\sqrt{2}$ .

(If you took the ratio of Ivan's to Anya's penny, then you'd get a ratio of  $1/2$  for final (kinetic) energy &  $1/\sqrt{2}$  for final speeds.)

#17



$$dU = -m_1 g dy_1 - m_2 g dy_2$$

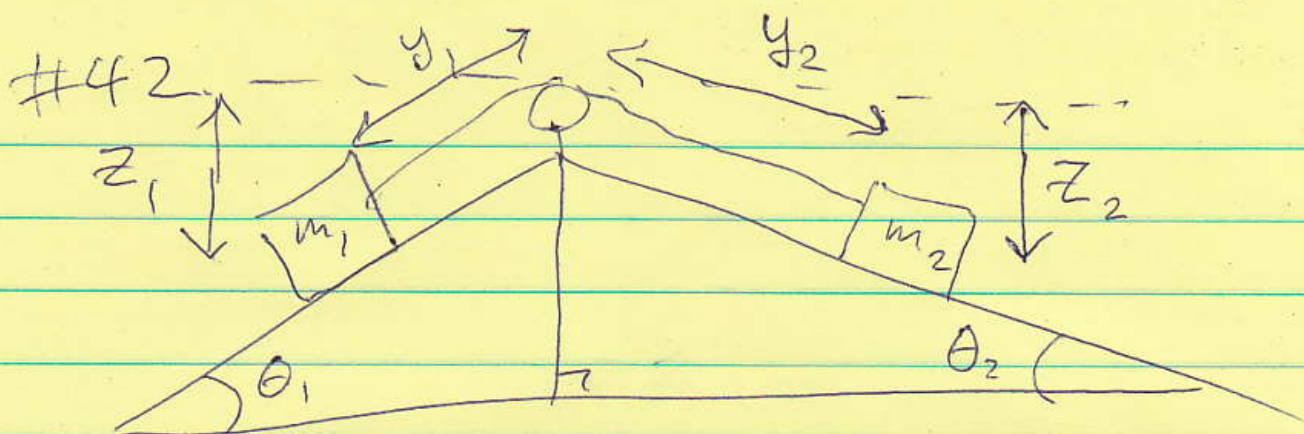
$$dy_2 = -2 dy_1$$

$$dU = (-m_1 + 2m_2) g dy_1$$

For balance, we need  $dU = 0$ ,

so we need  $-m_1 + 2m_2 = 0$ , so

$$m_1 / m_2 = 2$$



$$a_2 = \frac{dv_2}{dt} \quad v_2 = \frac{dy_2}{dt}$$

$$a_2 = ?$$

$$\Delta y_1 = -\Delta y_2$$

$$dy_1 = -dy_2$$

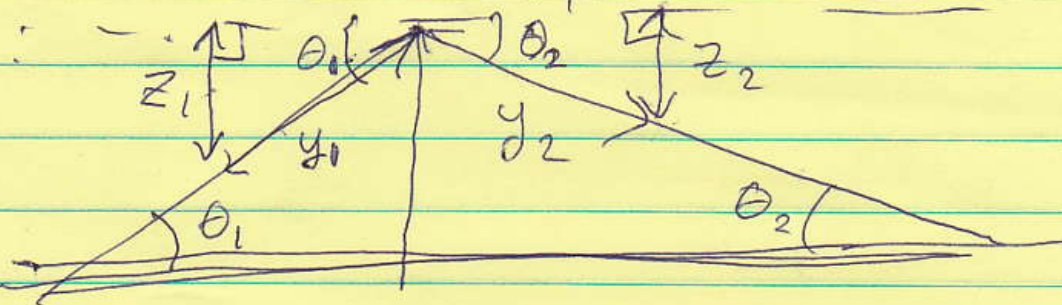
$$v_1 = \frac{dy_1}{dt} = -v_2$$

$$K = \frac{1}{2} m_1 v_1^2 + \frac{1}{2} m_2 v_2^2 = \frac{1}{2} v_2^2 (m_1 + m_2)$$

$(-v_2)^2 = v_2^2$

$$\Delta U = -m_1 g \Delta z_1 - m_2 g \Delta z_2$$

$$\sin \theta_1 = z_1 / y_1 \quad \sin \theta_2 = z_2 / y_2$$



$$z_1 = y_1 \sin \theta_1 \quad \Delta z_1 = \Delta y_1 \sin \theta_1$$

$$\Delta U = -m_1 g \Delta y_1 \sin \theta_1 - m_2 g \Delta y_2 \sin \theta_2$$

$$\Delta y_1 = -\Delta y_2$$

$$\Delta U = -m_1 g (-\Delta y_2) \sin \theta_1 - m_2 g \Delta y_2 \sin \theta_2$$

$$\Delta U = g \Delta y_2 (m_1 \sin \theta_1 - m_2 \sin \theta_2)$$

$$\frac{dU}{dt} = g \frac{dy_2}{dt} (m_1 \sin \theta_1 - m_2 \sin \theta_2)$$

$$dU/dt = g v_2 (m_1 \sin \theta_1 - m_2 \sin \theta_2)$$

$$dK + dU = 0$$

$$\frac{dK}{dt} + \frac{dU}{dt} = 0$$

Conservation of energy

$$K = \frac{1}{2} v_2^2 (m_1 + m_2)$$

$$dK = ?$$

$$dK = \frac{1}{2} (m_1 + m_2) d(v_2^2) = (m_1 + m_2) v_2 dv_2$$

$$2 v_2 dv_2$$

$$\frac{dK}{dt} = (m_1 + m_2) v_2 \frac{dv_2}{dt} = (m_1 + m_2) v_2 a_2$$

$$0 = \underbrace{\frac{dK}{dt} + \frac{dU}{dt}}_{=0} = (m_1 + m_2)v_2 a_2 + g v_2 (m_1 \sin \theta_1 - m_2 \sin \theta_2)$$

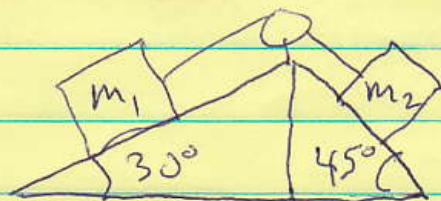
↑  
Solve for  $a_2$

Divide both sides by  $(m_1 + m_2)v_2$

$$0 = a_2 + \frac{g v_2 (m_1 \sin \theta_1 - m_2 \sin \theta_2)}{(m_1 + m_2)v_2}$$

$$\frac{g (m_2 \sin \theta_2 - m_1 \sin \theta_1)}{m_1 + m_2} = a_2$$

And  $a_1 = -a_2$



balanced if:

$$m_1 = \sqrt{2} m_2$$

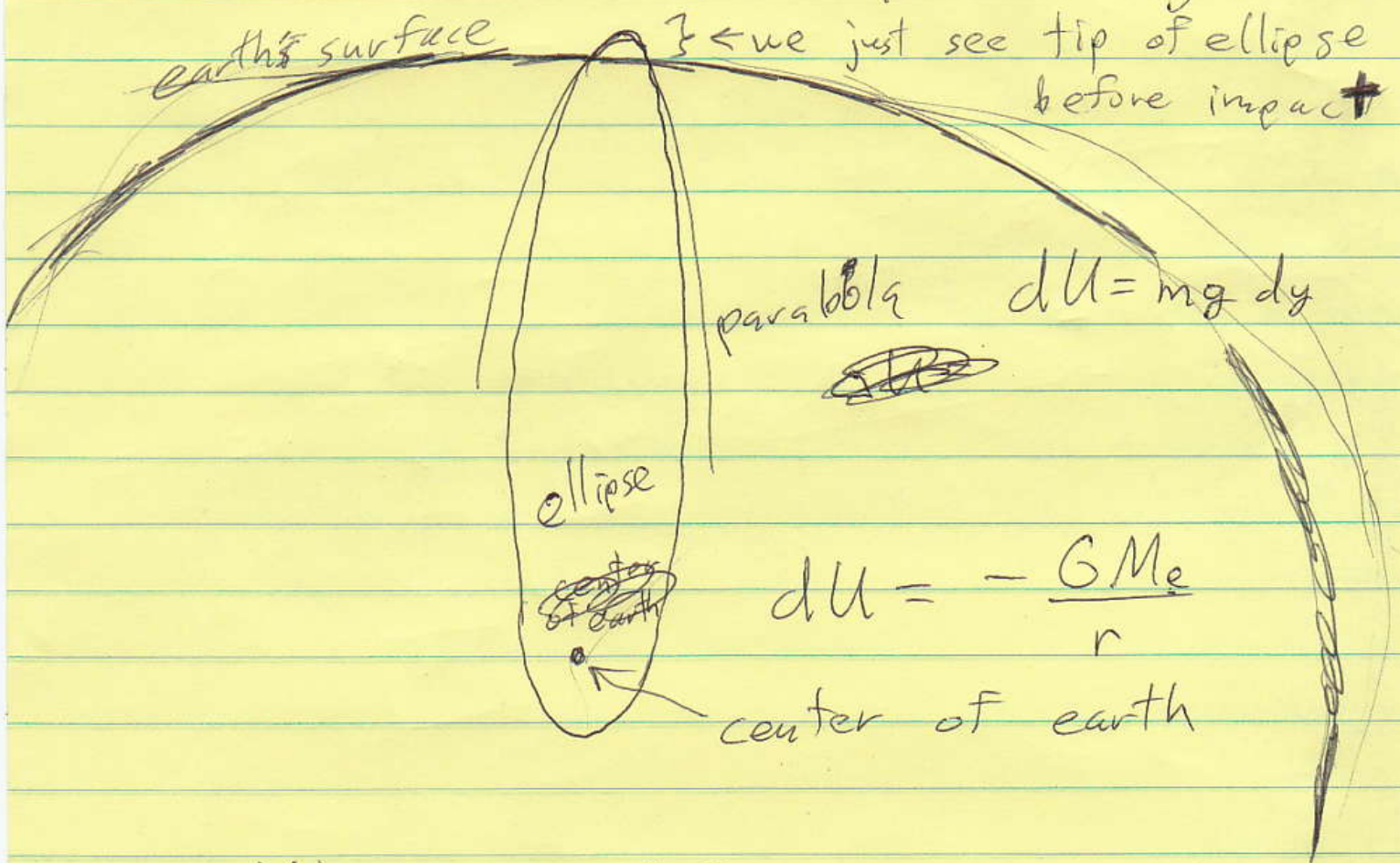
because

$$\sin 30^\circ = \frac{1}{\sqrt{2}} \sin 45^\circ$$



Back to universal gravitation:

Parabolic trajectories from gravity are just approximations of actual elliptical trajectories.



HW due 2/24: 21, 22, 24, 25, 27

Reading due 2/24: 2.4, 2.5